

IRAF

Redukcja w Długości Fali
i Strumieniach

Toma Tomov
CA UMK Torun

Final reduction steps

- **Wavelength calibration**
- **Flux calibration**

How to extract comparison spectra

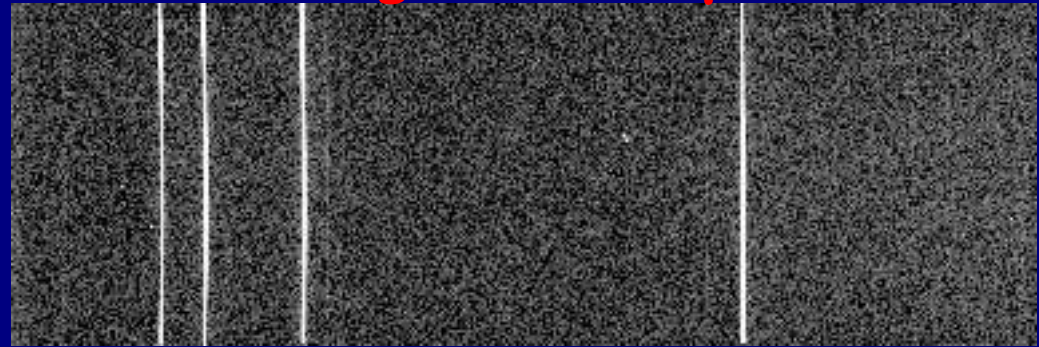
- First check that a line list for your comparison lamp exists in `/iraf/iraf/noao/lib/linelists/`. If there is not, you must prepare an ASCII file with wavelengths for your comparison spectrum.
- The comparison spectrum must be extracted exactly in like manner (the same pixels) as the star spectrum with a command similar to this

```
apall comp.imh ref=star.imh out=star_comp.imh \  
back- weight- clean-
```

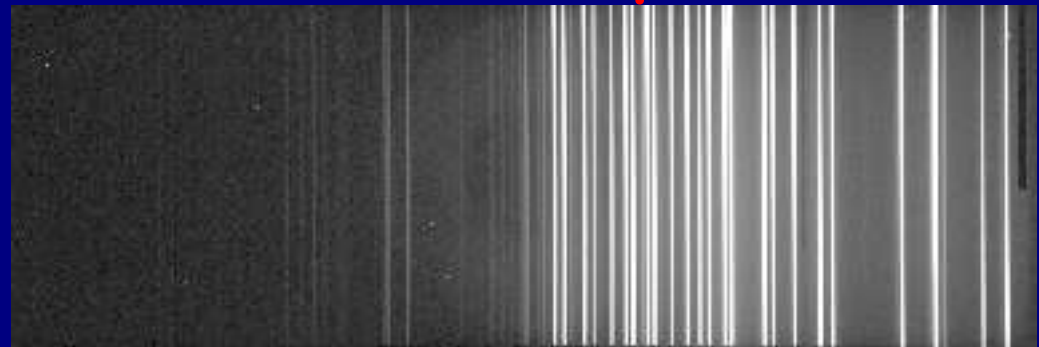
Our example in particular

- For a correct combine of the two lamp spectra we must:
 - process the images equally
 - avoid any pixels shift

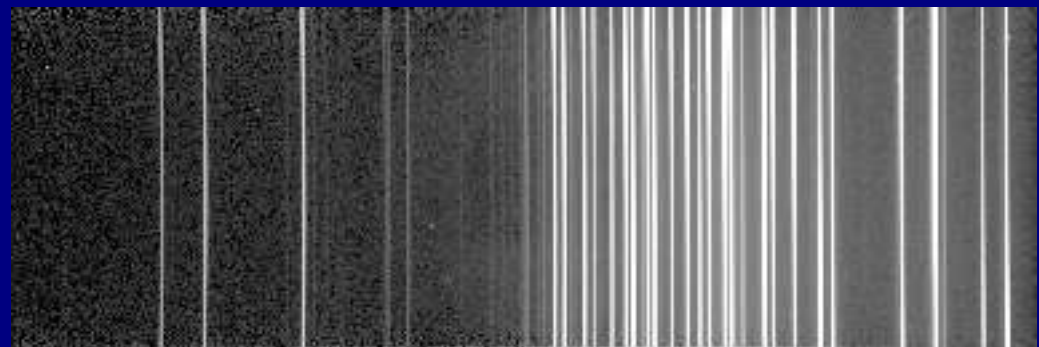
HgCdZn lamp



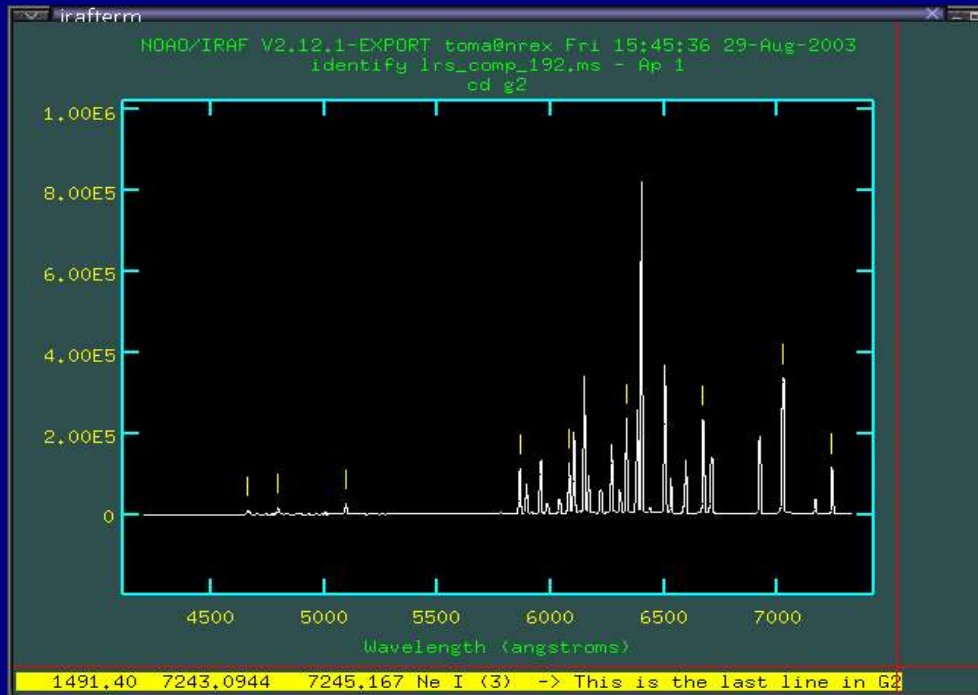
Ne lamp



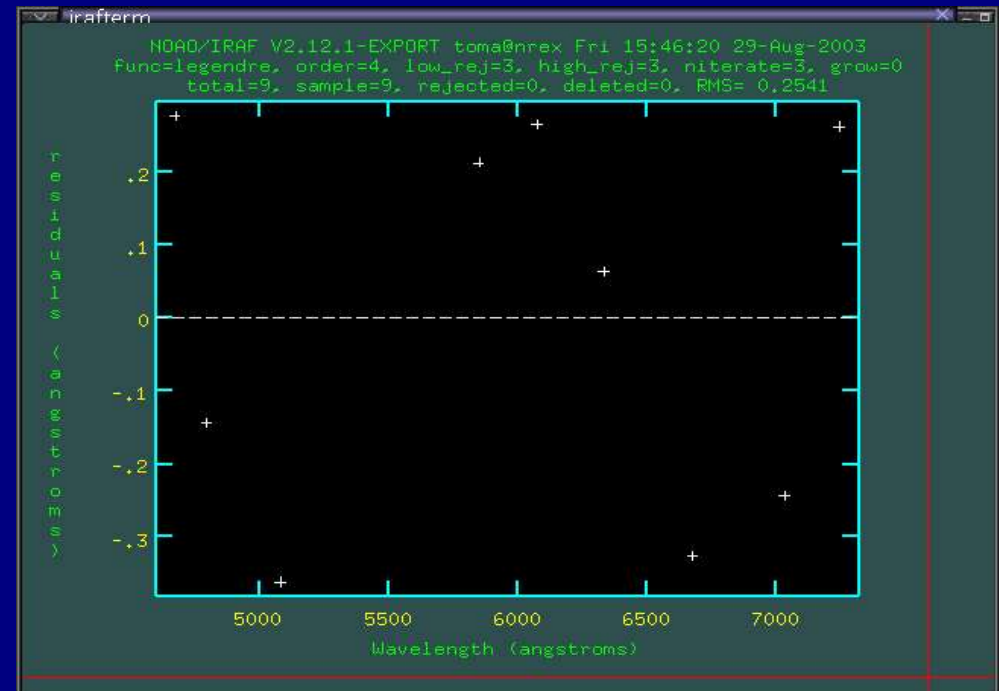
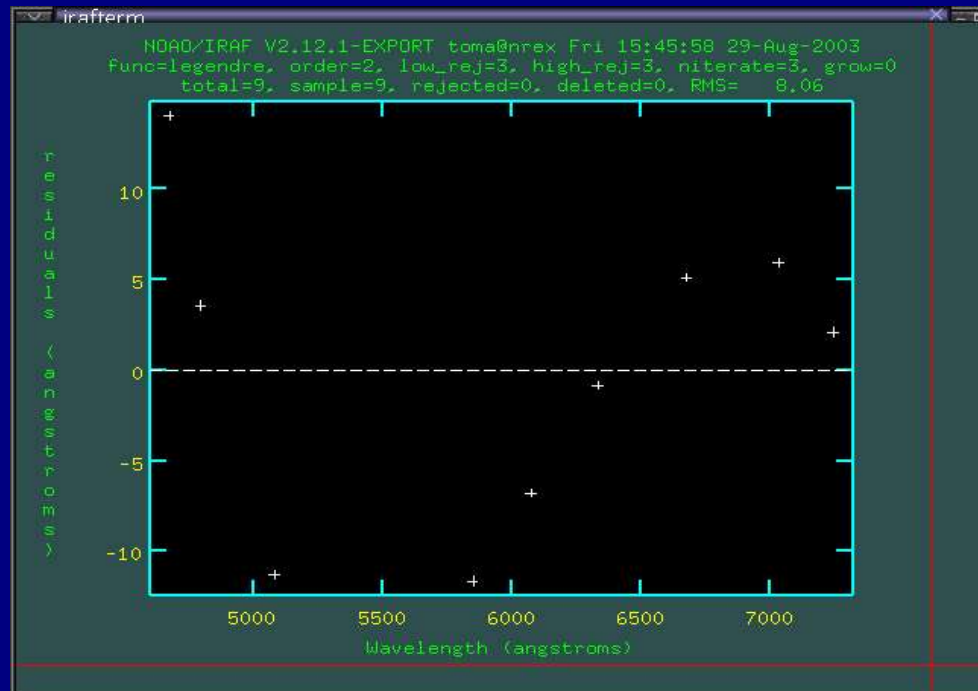
Sum



Identifying and fitting

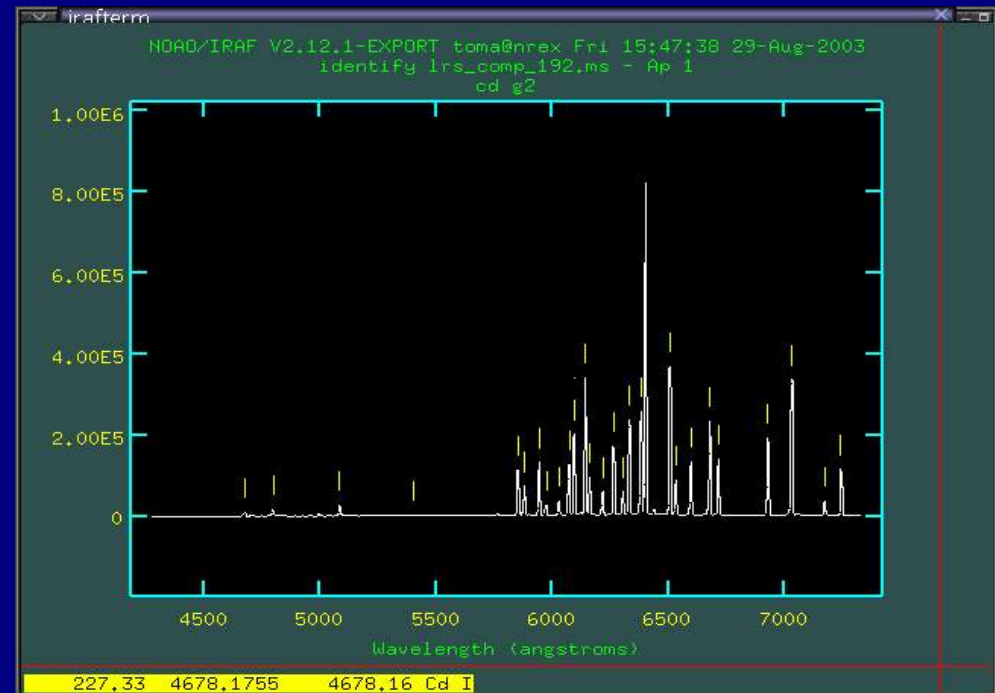


➤ **Initial dispersion solution (task identify)**

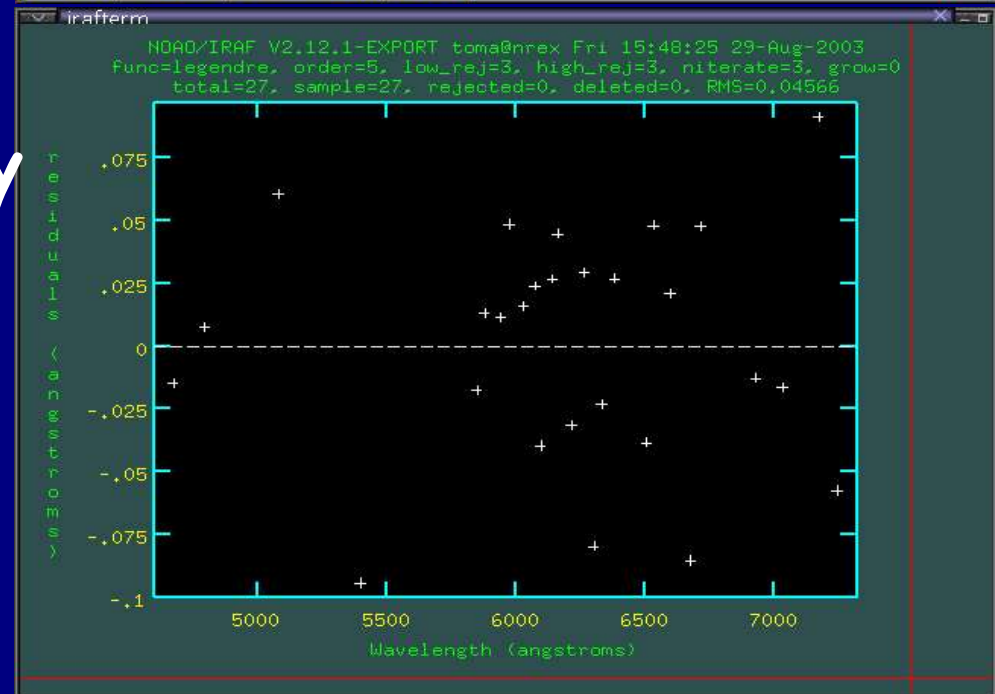


Identifying and fitting (cont.)

➤ **Final dispersion solution (task identify)**



➤ **Next step - use reidentify task for the rest of the comparison spectra obtained with the same set-up**



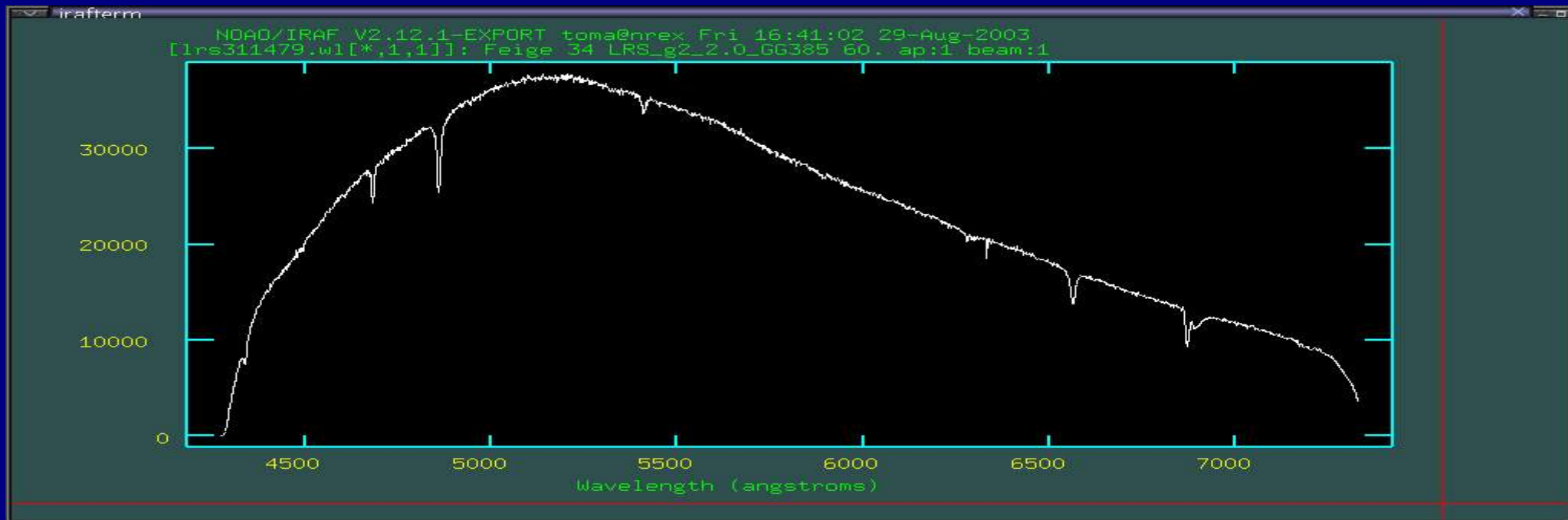
Apply wavelength calibration

- Tell to IRAF which wavelength solution to use

```
hedit star.ms.imh refspect1 comp.ms.imh
```

- Calibrate the science spectrum in wavelength

```
dispcor star.ms.imh star.wl.im linearize-
```

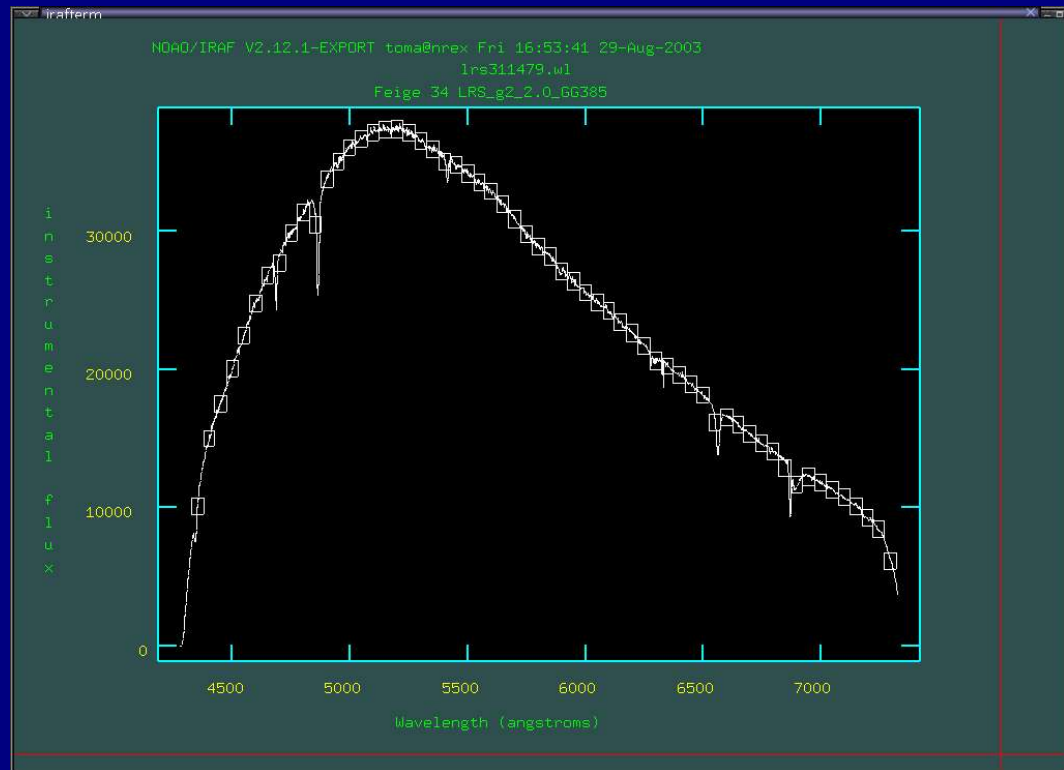


Flux calibration

➤ the task standard

- pay attention to have the needful info in the header
- pay attention to the extinction

```
lpar standard
  input = "lrs311479.wl"  Input image file root name
  output = "std"         Output flux file (used by SENSFUNC)
  star_name = "feige34"  Star name in calibration list
  airmass = 1.3          Airmass
  exptime = 300.         Exposure time (seconds)
  answer = "y"           (no|yes|NO|YES|NO!!|YES!)
  (samestar = yes)       Same star in all apertures?
  (beam_switch = no)     Beam switch spectra?
  (apertures = "")       Aperture selection list
  (bandwidth = INDEF)    Bandpass widths
  (bandsep = INDEF)      Bandpass separation
  (fnuzero = 3.6800000000000E-20) Absolute flux zero point
  (extinction = "onedstds$kpnoextinct.dat") Extinction file
  (caldir = "onedstds$spec50cal/") Directory containing calibration data
  (observatory = )_.observatory) Observatory for data
  (interact = yes)       Graphic interaction to define new bandpasses
  (graphics = "stdgraph") Graphics output device
  (cursor = "")          Graphics cursor input
  (mode = "ql")
```



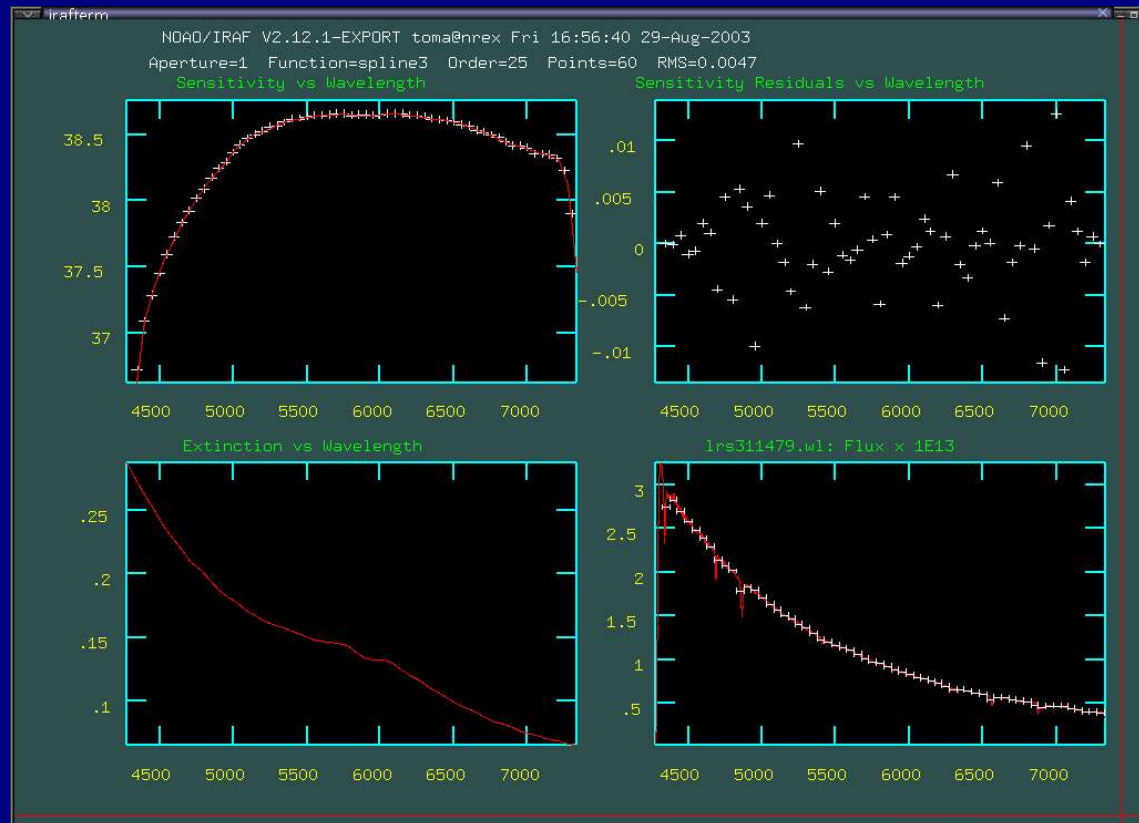
Flux calibration (cont.)

➤ the task sensfunc

- use high order to fit

- pay attention to the extinction

```
kp> lpar sensfunc
standards = "std"           Input standard star data file (from STANDARD)
sensitivity = "sens"       Output root sensitivity function imagename
                           (nolyes!NOIYES)
  answer = "yes"           Aperture selection list
  (apertures = " ")        Ignore apertures and make one sensitivity funct
  (ignoreaps = yes)        Output log for statistics information
  (logfile = "logfile")    Extinction file
  (extinction = "onedstds$/kpnoextinct.dat") Output revised extinction file
  (newextinctio = "extinct.dat") Observatory of data
  (observatory = )_observatory Fitting function
  (function = "spline3")    Order of fit
  (order = 6)              Determine sensitivity function interactively?
  (interactive = yes)      Graphs per frame
  (graphs = "sr")          Data mark types (marks deleted added)
  (marks = "plus cross box") Colors (lines marks deleted added)
  (colors = "2 1 3 4")    Graphics cursor input
  (cursor = "")           Graphics output device
  (device = "stdgraph")   (mode = "ql")
```



Flux calibration (cont.)

➤ the task **calibrate** is
the final step

```
kp> lpar calibrate
      input = "lrs311479.wl"  Input spectra to calibrate
      output = "lrs311479.flx" Output calibrated spectra
      airmass = 1.3           Airmass
      exptime = 900.         Exposure time (seconds)
      (extinct = yes)        Apply extinction correction?
      (flux = yes)           Apply flux calibration?
      (extinction = "onedstds$/kпноextinct.dat") Extinction file
      (observatory = " ")    Observatory of observation
      (ignoreaps = yes)     Ignore aperture numbers in flux calibration?
      (sensitivity = "sens") Image root name for sensitivity spectra
      (fnu = no)            Create spectra having units of FNU?
      (mode = "ql")
```

